

ALMA MATER STUDIORUM Università di Bologna

marcella.brusa3@unibo.it

Office: 4S6 (4 floor, UE3 Navile)

High Energy Astrophysics

Presentation of the course AA 2023/2024

Second semester End of February / End of May

Prof. Marcella Brusa DIFA - Dipartimento di Fisica e Astronomia "Augusto Righi"

high energy astrophysics (M. Longair):

"I take this term to mean the astrophysics of **high energy processes** and their application in astrophysical and cosmological contexts"

"For example, we need to explain how **the massive black holes** present in the nuclei of active galaxies can be studied, how **charged particles are accelerated** to extremely high energies in astronomical environments, the origins of enormous **fluxes of high energy particles and magnetic fields** in active galaxies, the physical processes in the **interiors and environments of neutron stars**, the nature of the **dark matter**, the expected fluxes of **gravitational waves** in extreme astronomical environments, and so on".



More in general, High Energy Astrophysics studies:

OBJECTS WHERE GRAVITY IS VERY STRONG - COMPACT OBJECTS (e.g. White dwarves, Neutron stars, X-ray binaries, black holes, SMBHs/AGN)

MATERIAL MOVING VERY FAST - RELATIVISTIC (e.g. jets, supernovae)

PROCESSES 'VERY HOT' OR ENERGETIC (e.g. gas in clusters of galaxies, SNR)

THE UNIVERSE ITSELF (cosmology)

The study of such objects and processes covers a very wide range of physics and a large variety of different types of physical objects

high energy astrophysics (M. Longair):

"I take this term to mean the astrophysics of **high energy processes** and their application in astrophysical and cosmological contexts"

"For example, we need to explain how **the massive black holes** present in the nuclei of active galaxies can be studied, how **charged particles are accelerated** to extremely high energies in astronomical environments, the origins of enormous **fluxes of high energy particles and magnetic fields** in active galaxies, the physical processes in the **interiors and environments of neutron stars**, the nature of the **dark matter,** the expected fluxes of **gravitational waves** in extreme astronomical environments, and so on".

high energy astronomy/astrophysics (Wikipedia):

"the study of astronomical objects that release electromagnetic radiation of highly energetic wavelengths. It includes X-ray astronomy, gamma-ray astronomy, and extreme UV astronomy, as well as studies of neutrinos and cosmic rays. The physical study of these phenomena is referred to as high-energy astrophysics. Astronomical objects commonly studied in this field may include black holes, neutron stars, active galactic nuclei, supernovae, supernova remnants, and gamma ray bursts."



More in general, High Energy Astrophysics studies:

OBJECTS WHERE GRAVITY IS VERY STRONG - COMPACT OBJECTS (e.g. White dwarves, Neutron stars, X-ray binaries, black holes, SMBHs/AGN)

MATERIAL MOVING VERY FAST - RELATIVISTIC (e.g. jets, **supernovae**)

PROCESSES 'VERY HOT' OR ENERGETIC (e.g. gas in **clusters of galaxies**, SNR)

THE UNIVERSE ITSELF (cosmology) For certain classes of objects a **large** fraction of the emitted energy *is* in the high energy band of the electromagnetic spectrum

We will focus on high frequencies (X, gamma)



More in general, High Energy Astrophysics studies:

OBJECTS WHERE GRAVITY IS VERY STRONG - COMPACT OBJECTS (e.g. White dwarves, Neutron stars, X-ray binaries, black holes, SMBHs/AGN)

MATERIAL MOVING VERY FAST - RELATIVISTIC (e.g. jets, supernovae)

PROCESSES 'VERY HOT' OR ENERGETIC (e.g. gas in clusters of galaxies, SNR)

THE UNIVERSE ITSELF (cosmology)

Sometimes these phenomena also manifest at longer wavelengths

We will use multi-wavelength info to understand



High Energy Astrophysics

Radiation vs. Particles vs. GW (no neutrinos, cosmic rays, gravitational waves)

see "Astroparticle Physics"

"Gravitational Waves Astrophusics and Cosmology"

Theory:

physical processes and emission mechanisms that imply high energy release; thermal and non thermal phenomena

Observational perspective:

Focus on cosmic sources emitting at High frequencies in the E.M. spectrum (extreme UV, X-ray, Gamma ray) Multiwavelength (and multimessenger) connections

Instrumental perspective:

see "Astronomical Instrumentation"

how data can be obtained: instruments, telescopes, detectors focus on space missions (the 2000 revolution: Chandra, XMM, Fermi, eROSITA)



INTRODUCTION TO HIGH-ENERGY ASTROPHYSICS

- **Recap on emission mechanisms:** blackbody/bremsstrahlung/Synchrotron Scattering processes; Plasma physics; Line production; collisional and ionization equilibrium.
- History of the high-energy Astronomy. Excursus on the properties of the main X-ray/Gamma-ray telescopes, future perspectives

Detection techniques for high-energy photons: CCDs in X-rays. Wolter-type telescopes and coded-aperture masks. Cerenkov radiation.









Stereoscopic imaging technique





ALMA MATER STUDIORUM Università di Bologna

ACCRETION PHYSICS:

- **Theory:** Eddington limit, accretion discs. Accretion and ejection processes (jets, winds).
- **Physics of compact objects**: white dwarfs, neutron stars, Chandrasekhar limit, pulsars, black holes (with some mentions on relativistic astrophysics)





GALACTIC HIGH-ENERGY SOURCES:

- X-ray emission from Solar System objects
- X-ray Emission from stars: main-sequence (MS) and pre-MS stars/very low-mass stars; the Sun
- Low- and High-mass X-ray binaries: phenomenology, classification and physical properties. Candidates black holes.
- Supernova Remnants



http://chandra.harvard.edu/photo/2004/casa/

http://www.lmsal.com/SXT/ .

http://chandra.harvard.edu/photo/2000/0065/

MW CONNECTION (1): SNR, The Crab Nebula





EXTRA-GALACTIC HIGH-ENERGY SOURCES:

- Clusters of galaxies and cooling flows
- Active Galactic Nuclei (AGN): X-ray spectral components: models vs. observations.
- X-ray surveys: obscured AGN contribution to cosmic X-ray background and feedback
- Ultraluminous X-ray sources
- Gamma ray bursts: phenomenologies, origin and emission mechanisms



XMM-Newton onbservations of the COSMOS field

MW/Multimessenger CONNECTION (2): The Kilonova and GW detection **MW CONNECTION (3)**: The Galactic center

Energy (keV) 6



Organization of Lectures

THEORY (~7)

TECHNOLOGIES (~3)

HE EMISSION FROM CLASSES OF SOURCES (SEMINAL AND MOST RECENT RESULTS) (~7) + MULTIWAVELENGTH CONNECTIONS (~3)

SPECIALISTIC SEMINARS (~4)

(by colleagues of OAS/CTA on future directions in HEA)



Textbooks

Main textbook:

Other textbooks:

Malcolm S. Longair: "*High-Energy Astrophysics*" (2011) Cambridge University Press High Energy Astrophysics

Malcolm S. Longair

Frank, A. King & D. Raine: "*Accretion Power in Astrophysics*" Cambridge University Press

George B. Rybicky, Alan P. Lightman: "*Radiative Processes in Astrophysics*" Wiley

Camenzind, Max: "Compact Objects in Astrophysics"

Freedman, R. & Kauffmann W: "Universe"

Prof. Marcella Brusa - DIFA / Intro

Other Resources:

papers and review articles on specific topics to discuss most recent advances and results from state-of-the art observations



Material

Main textbook:

Malcolm S. Longair: "*High-Energy Astrophysics*" (2011) Cambridge University Press High Energy Astrophysics

Malcolm S. Longair

Final thoughts and assignments

Longair, "High-Energy Astrophysics" (2011)

Read Chapter 6 (Section 6.2.3 fac.) Read Chapter 7.1 (all the rest fac.) Read Chapter 8.1-8.8 (Section 8.4 only the results 8.4.4; Sect. 8.7.2 fac.)

Refresh "The radio emission of the Galaxy" (Chapt. 8.9)

<u>Optional</u>: Read derivation of Larmor formula from classic electromagnetism (6.2.3)

Prof. Marcella Brusa - DIFA / Lecture 2



Final thoughts and assignments

Longair, "High-Energy Astrophysics" (2011) Read Chapters 14.7-14.8

Read Remillard & McClintock 2006 review on Black Hole binaries

Read Miller-Jones et al. 2021 (<u>https://arxiv.org/abs/2102.09091</u>) and understand how the new value of BH mass for Cy X-1 has been calculated

Homework:

- research how many XRBs in galaxies other than the Milky Way
- update the numbers of known BH binaries

Prof. Marcella Brusa - DIFA / Lecture 12

ALMA MATER STUDIORUM UNIVERSITÀ DI BOLOGNA



Questions?

ALMA MATER STUDIORUM UNIVERSITÀ DI BOLOGNA