RESOLVED STELLAR POPULATIONS Barbara Lanzoni A.A. 2023/2024



STELLAR ASTROPHYSICS

structure and physics of stars

STELLAR EVOLUTION

time variations of stellar structure



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observational properties of star clusters

and comparison with theory

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stars are individually resolved

for every single star:

- magnitude (colour)
- chemical composition
- radial velocity
- rotational velocity

from the ensemble of these individual stars:

- distance
- age
- kinematics

local dwarf galaxies





open clusters







check of Stellar Evolution theory and interpretation of non-resolved galaxies



and **ALSO**:

- estimate of age, distance, metallicity (from photometry)
- determination of the Spectral Energy Distribution (SED)

=> interpretation of distant (non-resolved) galaxies

Example:



tracers of the Milky Way formation history (=> of galaxies, in general)

ω Centauri in the MW halo



Terzan 5 in the MW bulge





tracers of the Milky Way formation history (=> of galaxies, in general)

$\boldsymbol{\omega}$ Centauri in the MW halo



remnant of a dwarf galaxy accreted by the MW?

Terzan 5 in the MW bulge



remnant of a structure that contributed to build the Galactic bulge?

study of multi-body dynamics and its effects on stellar evolution



tidal captures

collapse

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- discussion of both consolidated results, and ongoing research projects
- material: copy of the slides, scientific papers
- + oral exam
- Master Degree theses within the "stellar group" of DIFA (BL, Ferraro, Mucciarelli, Pallanca, Cadelano, Lardo; http://www.cosmic-lab.eu), and/or with INAF researchers (Origlia, Dalessandro, Bellazzini, Annibali, Romano,...)